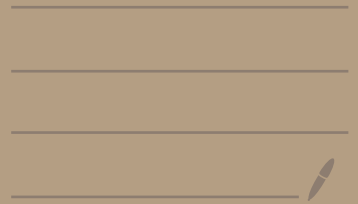


Classical Field Theory - Charles Torre



A mechanical system is a dynamical system w/ finitely many degrees of freedom.

A field is also a dynamical system but w/ infinitely many degrees of freedom.

Mathematically, a field is a section of a fiber bundle.

Let $\varphi: \mathbb{R}^4 \rightarrow \mathbb{R}$ be a scalar field
 $x^\alpha = (t, x, y, z)$

The Klein-Gordon Equation is

$$\square \varphi - m^2 \varphi = 0$$

where $\square = -\partial_t^2 + \partial_x^2 + \partial_y^2 + \partial_z^2$

a wave operator called the d'Alembertian

When $m=0$, we get the wave eqn:

$$\Delta \psi = \partial_t^2 \psi$$

↑
Laplacian

The KG eqn came about as an attempt to give a relativistic Schrödinger eqn but this did not work; after all, Schrödinger is mechanical, i.e. finitely many deg of freedom

But KG is sort of a classical limit of a quantum field in relativistic settings

How to solve KG? Suppose ψ is well behaved, such as $\psi_t \in L^2(\mathbb{R}^3) \forall t$.

Taking a Fourier expansion:

$$\psi(t, r) = \left(\frac{1}{2\pi}\right)^{3/2} \int_{\mathbb{R}^3} \hat{\psi}_k(t) e^{ik \cdot r} d^3k$$

We have $\hat{\psi}_{-k} = \hat{\psi}_k^*$ since ψ is real valued.

If ψ satisfies $\square\psi = m^2\psi$, then

$$\int -\partial_t^2 (\hat{\psi}_k(t)) e^{ik \cdot r} d^3r$$

$$+ \int \partial_x^2 (\hat{\psi}_k e^{ik \cdot r}) + \partial_y^2 (\hat{\psi}_k e^{ik \cdot r}) + \partial_z^2 (\hat{\psi}_k e^{ik \cdot r}) d^3r$$

$$= m^2 \int \hat{\psi}_k \cdot e^{ik \cdot r} d^3r$$

$$\begin{aligned} & \hat{\psi}_k \cdot \partial_z^2 (e^{ik \cdot r}) \\ &= -k_z^2 \hat{\psi}_k e^{ik \cdot r} \end{aligned}$$

\Rightarrow In the integrands:

$$\left(-\ddot{\hat{\psi}}_k - \underset{\substack{\uparrow \\ \text{dot product}}}{\vec{k} \cdot \vec{k}} \hat{\psi}_k \right) e^{ik \cdot r} = m^2 \hat{\psi}_k e^{ik \cdot r}$$

$$\Rightarrow \ddot{\hat{\psi}}_k + (k^2 + m^2) \hat{\psi}_k = 0.$$

\uparrow norm square



This is easy to solve; it's just a 2nd order ODE of a simple form.

$$\hat{\varphi}_k(t) = a_k e^{i\omega_k t} + b_k e^{-i\omega_k t}$$

$$\omega_k = \sqrt{k^2 + m^2}$$

Since $\hat{\varphi}_{-k} = \hat{\varphi}_k^*$, then $b_{-k} = a_k^*$.

KG field

$$\varphi(x) = \left(\frac{1}{2\pi}\right)^3 \int d^3k \left(a_k e^{i(k \cdot x - \omega_k t)} + a_k^* e^{-i(k \cdot x - \omega_k t)} \right)$$

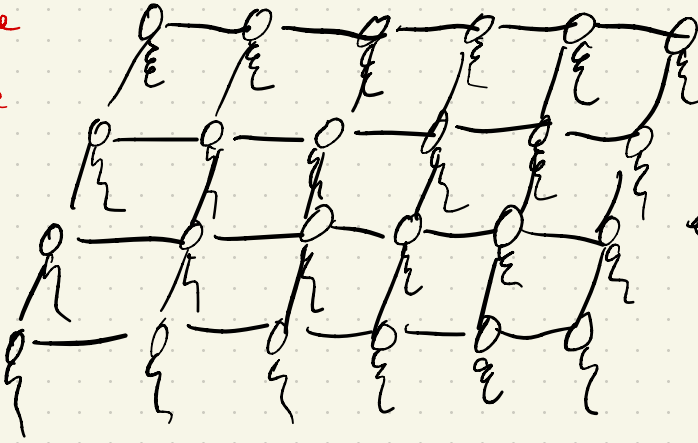
Note: The solution φ to KG is essentially an infinite collection of uncoupled harmonic oscillators for each $k \in \mathbb{R}^3$

So, there are infinite deg of freedom

Also, the KG field is often called the free or non-interacting field $\frac{1}{2}$ of the uncoupled nature of the oscillators

So picture an array of springs as a visual of a field

Discrete
picture



\mathcal{Q} satisfying
 $\mathbb{R}G$

← nudge it here
to "excite" the
array. Then
a wave will
propagate outwards.

This excitation which leads to a wave is what
we call a particle. Particles are waves in QM.

Let $L = \frac{1}{2} \int_{\mathbb{R}^3} (\dot{\varphi}^2 - |\nabla\varphi|^2 - m^2\varphi^2) d^3x$ (Lagrangian) on space
 $= \mathcal{L}$ (Lagrangian density)

$S[\varphi] = \int_{t_1}^{t_2} L dt$, (action), let $\mathcal{R} = [t_1, t_2] \times \mathbb{R}^3$
 include time now

Variation of S .

Let $\lambda \in \mathbb{R}$ be a parameter } φ_λ a 1-parameter family of fields.
 $\omega \varphi_0 = \varphi$

$\delta S = \frac{dS[\varphi_\lambda]}{d\lambda} \Big|_{\lambda=0}$. Also $\delta\varphi = \frac{d\varphi_\lambda}{d\lambda} \Big|_{\lambda=0}$,

Then $\delta S = \int_{\mathcal{R}} (\varphi \delta\varphi - \nabla\varphi \cdot \nabla\delta\varphi - m^2\varphi\delta\varphi) d^4x$.

Note: $\nabla \cdot (\nabla\varphi \cdot \delta\varphi) = \nabla^2\varphi \delta\varphi + \nabla\varphi \cdot \nabla\delta\varphi$.

Using integration by parts, note:

$$\int \ddot{\varphi} \delta\varphi = \dot{\varphi} \delta\varphi - \int \dot{\varphi} \delta\dot{\varphi}.$$

$$\delta S = \int_{\mathcal{R}} (-\ddot{\varphi} + \nabla^2 \varphi - m^2 \varphi) \delta\varphi d^4x + \left[\int_{\mathbb{R}^3} \dot{\varphi} \delta\varphi \right]_{t_1}^{t_2}$$

$$- \int_{\mathcal{R}} \nabla \cdot (\nabla \varphi \delta\varphi) d^4x.$$

$$= - \int_{t_1}^{t_2} dt \int_{r \rightarrow \infty} n \cdot \nabla \varphi \delta\varphi d^3A \quad (\text{Divergence Theorem})$$

If φ has cpt support or $\rightarrow 0$ faster than $\frac{1}{r^2}$, then the last term (boundary term) vanishes.

Other boundary conditions, such as the endpoints of φ are fixed will make $\delta\varphi|_{t_1} = \delta\varphi|_{t_2} = 0$.

This will force the middle term to vanish.

∴ w/ all these conditions,

$$\delta S = 0 \text{ when } \int_{\mathcal{R}} (\ddot{\varphi} + \nabla^2 \varphi - m^2 \varphi) \delta \varphi \, d^4 x = 0$$

$$\Rightarrow \square \varphi - m^2 \varphi = 0 \text{ everywhere in } \mathcal{R}.$$

So S is the correct action for KG eqn

Another view is via Euler-Lagrange eqn.

$$\mathcal{L} = \frac{1}{2} (\dot{\varphi}^2 - |\nabla \varphi|^2 - m^2 \varphi^2)$$

Treating $x, \varphi, \varphi_\alpha$ as formal variables (so \mathcal{L} is in the 1^{st} Jet space)

$$\text{define } E(\mathcal{L}) = \frac{\partial \mathcal{L}}{\partial \varphi} - \underbrace{D_\alpha \frac{\partial \mathcal{L}}{\partial \varphi_\alpha}}_{\text{total derivative}}$$

Then $\frac{\delta S}{\delta \varphi} = \mathcal{E}(\mathcal{L}) \Big|_{\varphi=\varphi(x)} = \square \varphi - m^2 \varphi.$

In fact, if $\hat{\mathcal{L}} = \frac{1}{2} \varphi (\square - m^2 \varphi)$, $\mathcal{E} \hat{\mathcal{L}}^2$ (2nd jet space) one can show

$\mathcal{E}(\mathcal{L}) = \mathcal{E}(\hat{\mathcal{L}}) \quad \{ \quad \mathcal{L} = \hat{\mathcal{L}} + \text{divergence term}$
sum of all the 1st partial derivatives of something.

Some Generalizations of KG.

let $\mathcal{L} = \frac{1}{2} (\varphi^2 - |\nabla \varphi|^2 - m^2 \varphi^2) - j \varphi$

where $j: \mathbb{M}^4 \rightarrow \mathbb{R}$ is the source (in electro magnetism, j could be electric charge or current)

Then $\mathcal{E}(\mathcal{L}) = (\square - m^2) \varphi - j.$

In QFT, the presence of a source leads to particle creation/annihilation via transfer of energy-momentum b/w the field & its source.

The KG eqns are linear \therefore so the solutions are non-interacting.

If we modify KG to be non-linear, we introduce self-interaction.

$$\text{Eg. } \mathcal{L} = \frac{1}{2} (\dot{\varphi}^2 - |\nabla\varphi|^2 - m^2\varphi^2) - V(\varphi)$$

$$E(L) = (\square - m^2)\varphi - V'(\varphi) = 0.$$

So long as V isn't quadratic, this becomes non-linear.

One potential of interest: $V(\varphi) = -\frac{1}{2} a^2 \varphi^2 + \frac{1}{4} b^2 \varphi^4$.

If V is quad, then $V'(\varphi) = a\varphi - b$

$$\therefore \text{so } (\square - (m^2 + a))\varphi = b$$

Seems like, you just increase the mass, — what does this mean, physically?

Coordinate free description: let (M, g) be a Lorentzian mfd.

$$\text{Then } \mathcal{L} = -\frac{1}{2} (g^{-1}(d\varphi, d\varphi) + m^2 \varphi^2) \varepsilon(g)$$

volume form
of g .

is the Lagrangian density.

This is called minimally coupled

Let ξ be a parameter ξ $R(g) = \text{scalar curvature}$. Then

$$\mathcal{L} = -\frac{1}{2} [g^{-1}(d\varphi, d\varphi) + (m^2 + \xi R(g)) \varphi^2] \varepsilon(g) \text{ gives}$$

Curvature coupled KG theory.

These theories are not "diffeomorphism invariant" or
"generally covariant"; i.e. $f: M \xrightarrow{\text{diffeo}} M$ $\} \hat{g} = f^* g$, then

$$\hat{\mathcal{L}} = -\frac{1}{2} (\hat{g}^{-1}(d\varphi, d\varphi) + m^2 \varphi^2) \varepsilon(\hat{g}) \quad \text{is a new}$$

Lagrangian density in general unless f is a symmetry.

If we allow g to vary, we get 11 coupled nonlinear field eqns instead of 1 linear field eqn.

Conservation Laws are fundamental } give info about complicated dynamics. Also, conservation laws are related to symmetries by Noether's theorem

def: let $j^\alpha = j^\alpha(x, \varphi, \partial\varphi, \dots, \partial^k\varphi) \in \mathbb{R}^k$ be a vector field constructed as a local F^n . j^α is a conserved current or defines a conservation law if the divergence of $j^\alpha = 0$ when φ satisfies the field eqn.

eg. $D_\alpha j^\alpha = 0$ when $(\square - m^2)\varphi = 0$.

Explicitly, if $\varphi(x)$ is a solution, take

$$j^\alpha(x) \doteq j^\alpha(x, \varphi(x), \frac{\partial \varphi(x)}{\partial x}, \dots, \frac{\partial^k \varphi(x)}{\partial x^k}) \quad \text{s.t.}$$

$$\frac{\partial}{\partial x^\alpha} j^\alpha = 0.$$

If $j^\alpha = (j^0, j^1, j^2, j^3)$, call $\rho \doteq j^0$ the density

$\vec{j} = (j^1, j^2, j^3)$ the
current
density

$$\Rightarrow \frac{\partial \rho}{\partial t} + \nabla \cdot \vec{j} = 0.$$

The point of writing this is:

Let $Q_V(t) = \int_V \rho(t, \vec{x}) d^3x$ be the total charge in region V .

$$\text{Then } \frac{d}{dt} Q_V(t) = - \int_V \nabla \cdot \vec{j} = - \underbrace{\int_{\partial V} \vec{j} \cdot d\vec{S}}_{\text{net flux}}$$

We say Q_V is conserved since we can see how it changes over time by purely in terms of the ^{net} flux, a fixed value. So there's no creation nor destruction of charge; it just moves around.

If we place boundary conditions, such as $V = \mathbb{R}^3$; the field vanishes rapidly enough at ∞ , then $\frac{d}{dt} Q_V(t) = 0$; so the total charge is constant

Conservation of Energy

$$\text{let } \dot{j}^0 = \frac{1}{2} (\dot{\psi}^2 + |\nabla\psi|^2 + m^2\psi^2)$$

$$\dot{j}^i = -\dot{\psi}(\nabla\psi)_i$$

$$\left. \begin{aligned} \partial_0 \dot{j}^0 &= \dot{\psi}\ddot{\psi} + \nabla\psi \cdot \nabla\dot{\psi} + m^2\psi \cdot \dot{\psi} \\ \partial_i \dot{j}^i &= -\nabla\dot{\psi} \cdot \nabla\psi - \dot{\psi} \nabla^2\psi \end{aligned} \right\} \Rightarrow \partial_\alpha \dot{j}^\alpha = -\dot{\psi}(\square\psi - m^2\psi)$$

Sum the Einstein notation

So if ψ is a solution to KG, $\partial_\alpha j^\alpha = 0$.

Let $E_V = \frac{1}{2} \int_V (\dot{\psi}^2 + |\nabla\psi|^2 + m^2\psi^2) d^3x$

total energy

= T + U

$\frac{1}{2} \int_V \dot{\psi}^2 d^3x$

kinetic energy

$\frac{1}{2} \int_V (|\nabla\psi|^2 + m^2\psi^2) d^3x$

potential energy

Then $\rho = j^0$ $\vec{j} = (j^1 j^2 j^3)$ $\nabla_\mu (L \psi \nabla_\mu \psi)$

$$\frac{d}{dt} E_V = \int \partial_t \rho = - \int \nabla \cdot \vec{j} = - \int_V \nabla \psi \cdot \nabla \dot{\psi} + \dot{\psi} \nabla^2 \psi d^3x$$

$$= - \int_{\partial V} \dot{\psi} \nabla \psi dS$$

If $\dot{\psi} = 0$ or the normal component of $\nabla\psi$ to ∂V vanishes,

then $\frac{d}{dt} E_V = 0$. $\{$ Energy is conserved.